

# Considering Dispersive Interactions in the COSMO-SAC Model: A Study on Nonpolar Mixtures

Chieh-Ming Hsieh, Shiang-Tai Lin, Jadran Vrabec

## 1. Introduction

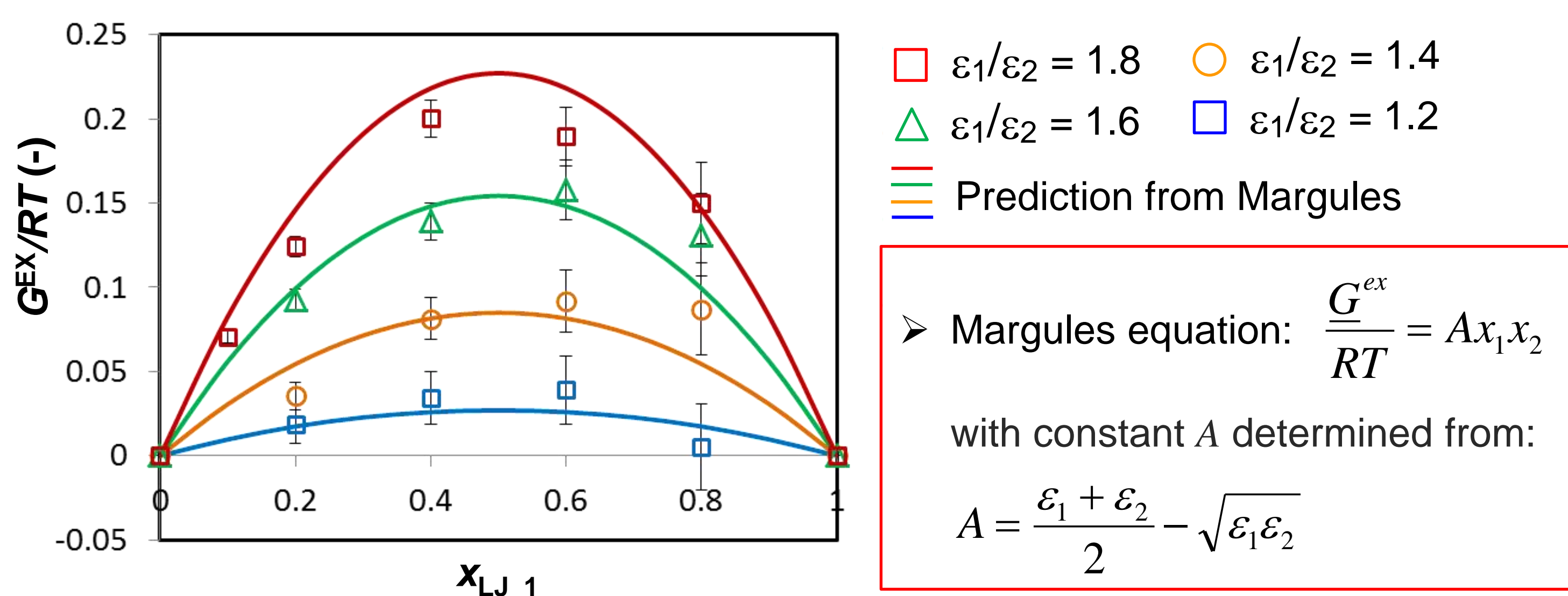
- A reliable predictive thermodynamic model is highly desirable to reduce the amount of required experimental data, which are crucial for the design of chemical and pharmaceutical processes.
- The COSMO-SAC model, that is exclusively based on the results from quantum chemical calculations, does not contain any species-dependent parameters and thus never suffers with respect to missing parameters. However, the dispersive intermolecular interactions in mixtures are neglected. This is considered as the main reason why the COSMO-SAC model fails to yield good predictions for nonpolar mixtures.
- In this study, we propose a new dispersion term to explicitly consider the contribution of the dispersive interactions to the non-ideality of mixtures from classical force field based molecular simulation results.

## 2. Theory

### A. Studying Dispersive Interactions with Molecular Simulation

#### Mixtures of Lennard-Jones fluids

	$\epsilon_1$	$\epsilon_2$	$\epsilon_1/\epsilon_2$	$\sigma_1/\sigma_2$	Margules A
LJ_1	100	83.333	1.2	1	0.102
LJ_2		71.428	1.4	1	0.324
		62.500	1.6	1	0.592
		55.556	1.8	1	0.875



Since the single constant Margules equation can well describe  $G^{EX}$  for systems containing only dispersive interactions, it was used here to consider the dispersive contribution to the non-ideality of mixtures in COSMO-SAC model.

### B. COSMO-SAC-dsp Model

$$\ln \gamma_{i/S} = \frac{\Delta G_{i/S}^{*res} - \Delta G_{i/i}^{*res}}{RT} + \ln \gamma_{i/S}^{SG,comb} + \ln \gamma_{i/S}^{disp}$$

Residual contribution: Electrostatic interactions between molecules in the mixture

$$\frac{\Delta G_{i/S}^{*res}}{RT} = n_i \sum_{\sigma_m} p_i(\sigma_m) \ln \Gamma_s(\sigma_m)$$

Using quantum chemistry results

Combinatorial contribution: Size and shape effect via Staverman-Guggenheim model:

$$\ln \gamma_{i/S}^{comb} = \ln \frac{\phi_i}{x_i} + \frac{z}{2} q_i \ln \frac{\theta_i}{\phi_i} + l_i - \frac{\phi_i}{x_i} \sum_j x_j l_j$$

$$\theta_i = \frac{x_i q_i}{\sum_j x_j q_j} \quad \phi_i = \frac{x_i r_i}{\sum_j x_j r_j} \quad l_i = \frac{z}{2} (r_i - q_i) - (r_i - 1)$$

$r_i$  and  $q_i$  are the normalized volume and surface area parameters for component  $i$

These two contributions were taken from previous work without any modification or parameter re-optimization.

#### Dispersion contribution (for binary mixtures $i + j$ )

$$\ln \gamma_{i/S}^{disp} = Ax_{j+i}^2 \quad \text{with constant } A \text{ determined from: } A = \frac{\epsilon_i + \epsilon_j}{2} - \sqrt{\epsilon_i \epsilon_j}$$

#### Dispersion parameter of a molecule

$$\epsilon_{Molecule} = \frac{1}{N_{Atom}} \sum_{j=1}^n \epsilon_j N_j$$

where:  $n$ : number of atoms in the molecule  
 $N_{atom}$ : number of atoms in the molecule  
 $N_j$ : number of type  $j$  atoms  
 $\epsilon_j$ : dispersion parameter of atom  $j$

#### Two different approaches were tested for systems containing compounds composed of C, H, N, O, F and Cl

- dsp1 (Npar=5,  $\epsilon_C$ ,  $\epsilon_O$ ,  $\epsilon_N$ ,  $\epsilon_F$  and  $\epsilon_{Cl}$ )
- dsp3 (Npar=10,  $\epsilon_{C(sp3)}$ ,  $\epsilon_{C(sp2)}$ ,  $\epsilon_{C(sp)}$ ,  $\epsilon_{O(=O)}$  and  $\epsilon_{O(-OH)}$ ,  $\epsilon_{O(=O)}$ ,  $\epsilon_{N(sp3)}$ ,  $\epsilon_{N(sp2)}$ ,  $\epsilon_{N(sp)}$ ,  $\epsilon_F$  and  $\epsilon_{Cl}$ )

## Acknowledgement

Financial Support from Alexander von Humboldt Foundation



## 3. Preliminary Results

### A. Overview on Vapor-Liquid Equilibrium Predictions

Comparing VLE predictions from COSMO-SAC-dsp, COSMO-SAC and modified UNIFAC

	VLE_nhb (Optimization)			VLE_hb (Prediction)		
	Nbinary	AARD-P	AAD-y1	Nbinary	AARD-P	AAD-y1
mUNIFAC(1998)	198	2.48	0.96	164	3.37	1.39
COSMO-SAC(2010)	221	4.66	1.59	170	6.22	2.17
COSMO-SAC-dsp1 (Npar: 5)	221	4.19	1.54	170	5.98	2.09
COSMO-SAC-dsp3 (Npar: 10)	221	3.84	1.47	170	6.07	2.14

PS. (1) VLE\_nhb: in total 7831 VLE data points; (2) VLE\_hb: in total 8598 VLE data points; (3) Different isotherms of the same binary mixture were considered as one binary; (4) Npar is the number of newly introduced dispersion parameters in the model

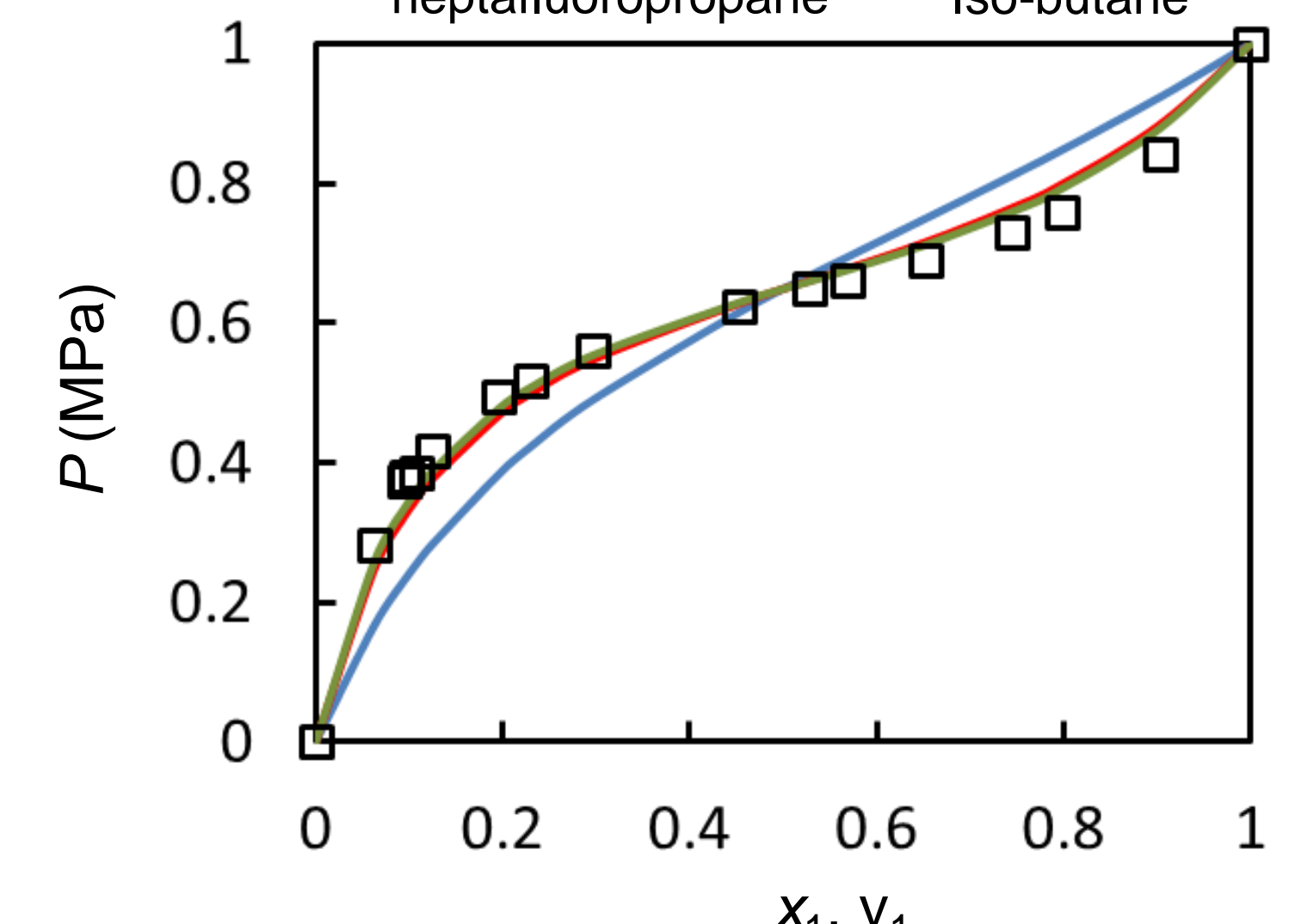
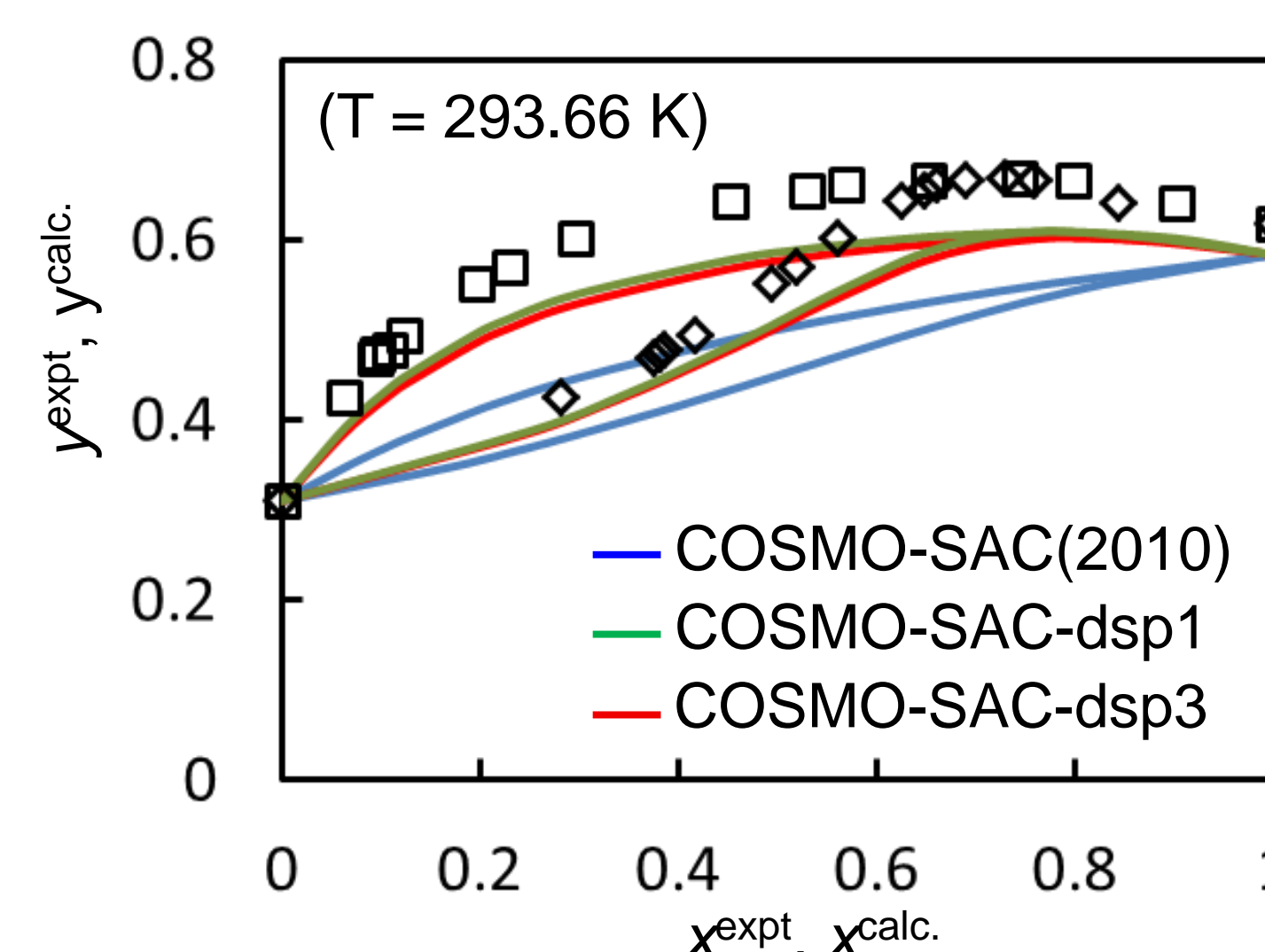
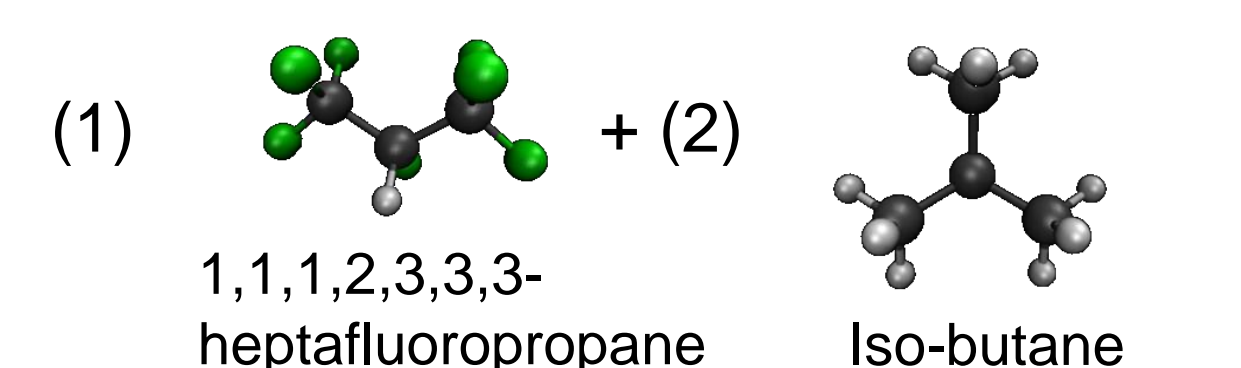
- A reduction of 15% in AARD-P can be achieved for non-hydrogen bonding systems!
- For hydrogen-bonding systems, dominated by electrostatic interactions, a reduction of 3% in AARD-P can be observed.

### B. Improvement of VLE prediction for systems containing fluoro-compounds

Comparing VLE predictions from COSMO-SAC-dsp, COSMO-SAC and modified UNIFAC

	VLE		
	Nbinary	AARD-P	AAD-y1
mUNIFAC(1998)	3	4.57	2.03
COSMO-SAC(2010)	16	11.22	3.97
COSMO-SAC-dsp1	16	6.53	3.65
COSMO-SAC-dsp3	16	6.88	3.70

- Modified UNIFAC has a severe issue of missing parameters for this type of system (only 3 out of 16 binary mixtures can be described).
- COSMO-SAC-dsp could be a good candidate for this type of mixtures.

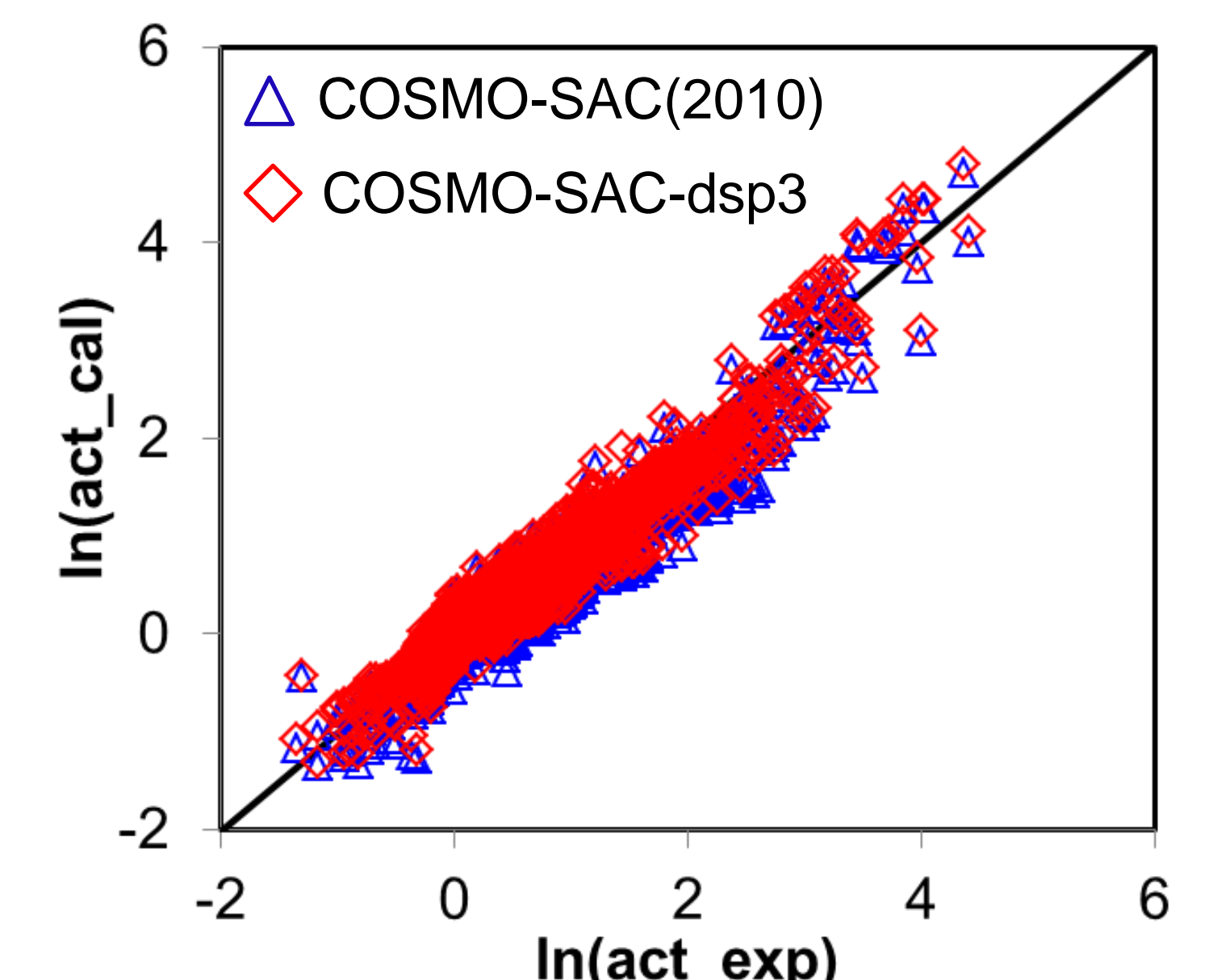


### C. Prediction of Infinite Dilution Activity Coefficient (IDAC)

Comparing prediction of IDAC from COSMO-SAC-dsp, COSMO-SAC and modified UNIFAC (in total 728 nhb binary mixtures)

	IDAC_nhb (Prediction)	
	Ndata	Error
mUNIFAC(1998)	1923	0.104
COSMO-SAC(2010)	1923	0.255
COSMO-SAC-dsp1	1923	0.219
COSMO-SAC-dsp3	1923	0.192

$$Error = \frac{1}{N} \sum_{j=1}^N |\ln \gamma_i^{\infty,expt} - \ln \gamma_i^{\infty,calc}|$$



- A reduction of 25% in terms of overall deviation can be achieved for non-hydrogen-bonding systems!

## 4. Conclusions

- A new dispersion contribution to the activity coefficient was proposed in the COSMO-SAC model.
- New COSMO-SAC-dsp model can provide better predictions, especially for non-hydrogen-bonding systems.

### References:

- C.-M. Hsieh; S. I. Sandler; S.-T. Lin, *Fluid Phase Equilib.*, 2010, 297, 90-97.
- S. Deublein; B. Eckl; J. Stoll; S. V. Lishchuk; G. Guevara-Carrion; C. W. Glass; T. Merker; M. Bernreuther; H. Hasse; J. Vrabec, *Comput. Phys. Commun.*, 2011, 182, 2350-2367.
- J. Gmehling, J. Lohmann, A. Jakob, J. Li, R. Joh, *Ind. Eng. Chem. Res.*, 1998, 37, 4876-4882.